



Active Antenna Design and Characterization for the mid-SKA

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SKA



- Memo100:
 - Sparse aperture array 70-450MHz
 - 4000 m²/K, T_{sys}~1000: 4 km²
 - 2 million Antennas
 - Dense aperture array 300-1000MHz
 - 10.000 m²/K, T_{sys}=50K: 0,9 km², nearly 1 km²
 - 45 million Antennas
 - Aperture efficiency 80%, 45 degree scan



Noise budget



Aperture Array	
Spill-over	0 K
Vivaldi feed loss	9 K
Low Noise Amplifier	16 K
Noise mismatch / coupling / 2 nd stage	7 K
Sky	8 K
Total	40 K

- $T_{LNA} \sim 30\%$ of noise budget only!



300K

- Cryogenic cooling not possible (€s!)
 - And not needed!?
- Just a wide band matching issue?!
 - Providing $f_T > 100\text{GHz}$
- mHEMT GaAs, CMOS

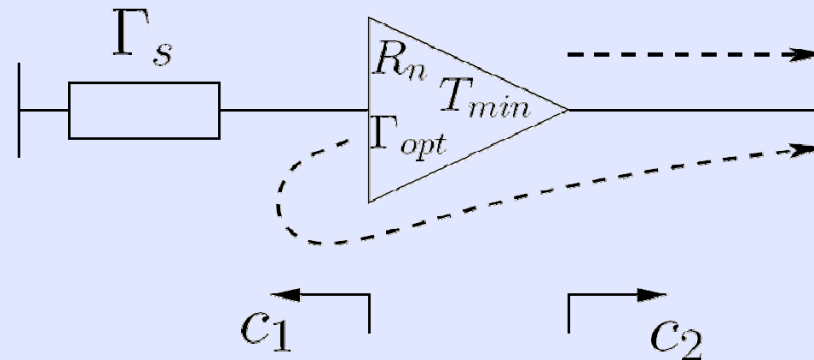
$$F_{MIN} = 1 + \frac{f}{f_T} \times \sqrt{g_m (r_g + r_s)}$$

$$Tn_{297} = \sqrt{20} Tn_{15}$$

$$T_{min} = \sqrt{T_{amb} T_{channel}} \frac{f}{f_T}$$



Aperture Array T_{sys}

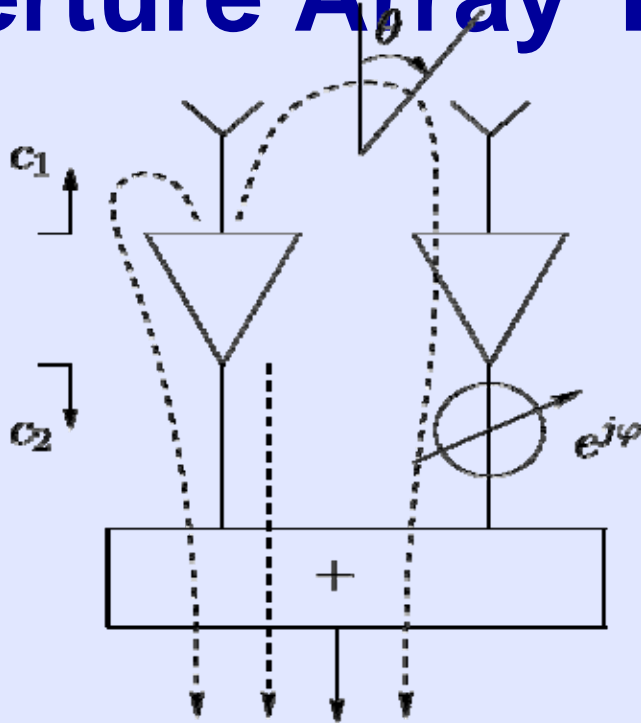


$$T_{LNA} = T_{min} + \frac{4T_0 R_n}{Z_0} \frac{|\Gamma_s - \Gamma_{opt}|^2}{|1 + \Gamma_{opt}|^2 (1 - |\Gamma_s|^2)}$$

- Single LNA case
- Noise Wave analyses
 - Optimum Noise match if $\Gamma_{opt} = \Gamma_s$



Aperture Array T_{sys}



$$T_{LNA} = T_{min} + \frac{4T_0 R_n}{Z_0} \frac{|\Gamma_s - \Gamma_{act}|^2}{|1 + \Gamma_{opt}|^2 (1 - |\Gamma_{act}|^2)}$$

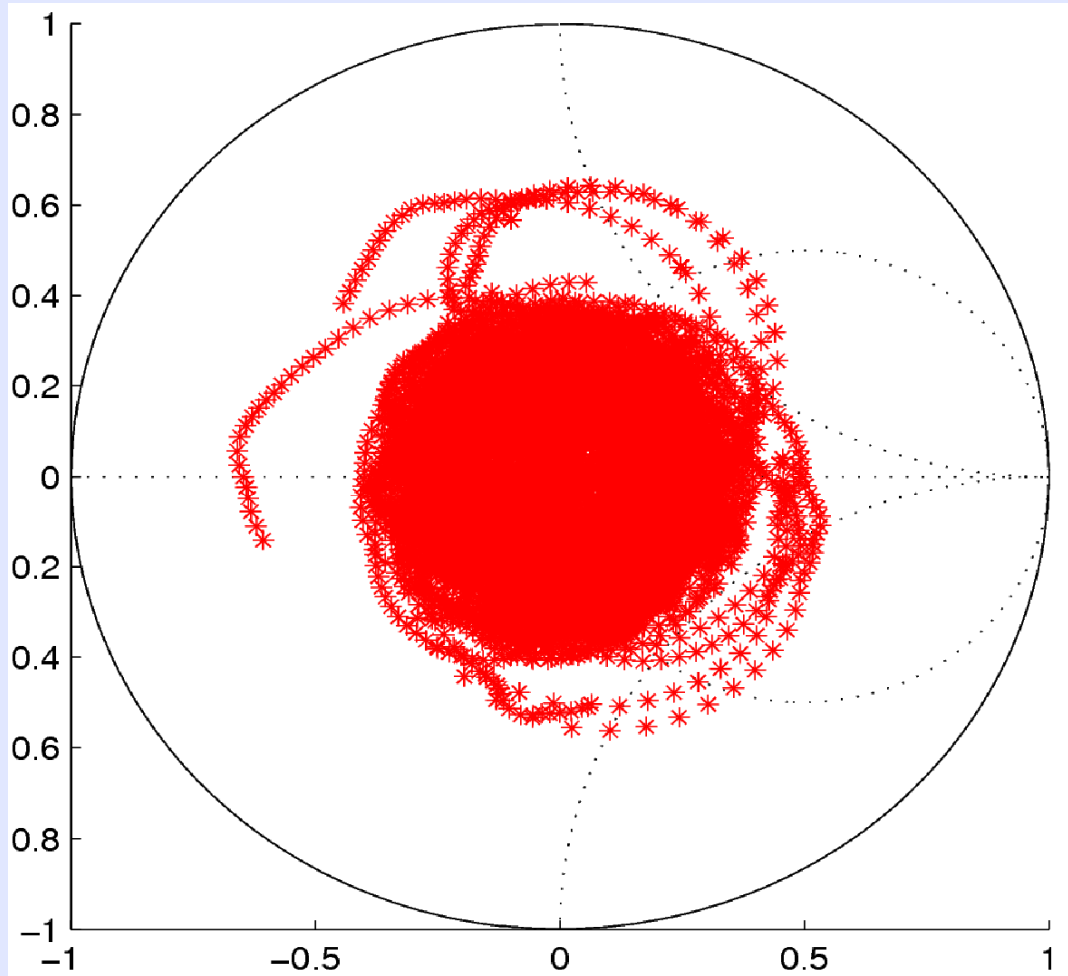
$$\Gamma_i(\theta, \varphi) = \frac{1}{a_i} \sum_{j=1}^N S_{ij} a_j$$

$$a_j = e^{-jk_0[(k-1)d_x \sin\theta \cos\varphi + (l-1)d_y \sin\theta \sin\varphi]}$$

- $T = \text{Direct part} + \text{Reflected part} + \text{coupled part}$
 - Optimum Noise match if $\Gamma_{opt} = \Gamma_{act}$



Active reflection coefficient while scanning



Measured

8×8 array

Γ_{act} of centre element

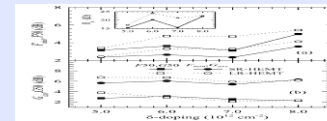
$500 < f < 1500$ MHz

$\theta < 45$ deg



Aperture Array T_{sys}

- Clearly low T_{min} and R_N or N required
- Limited work on N (process) optimization



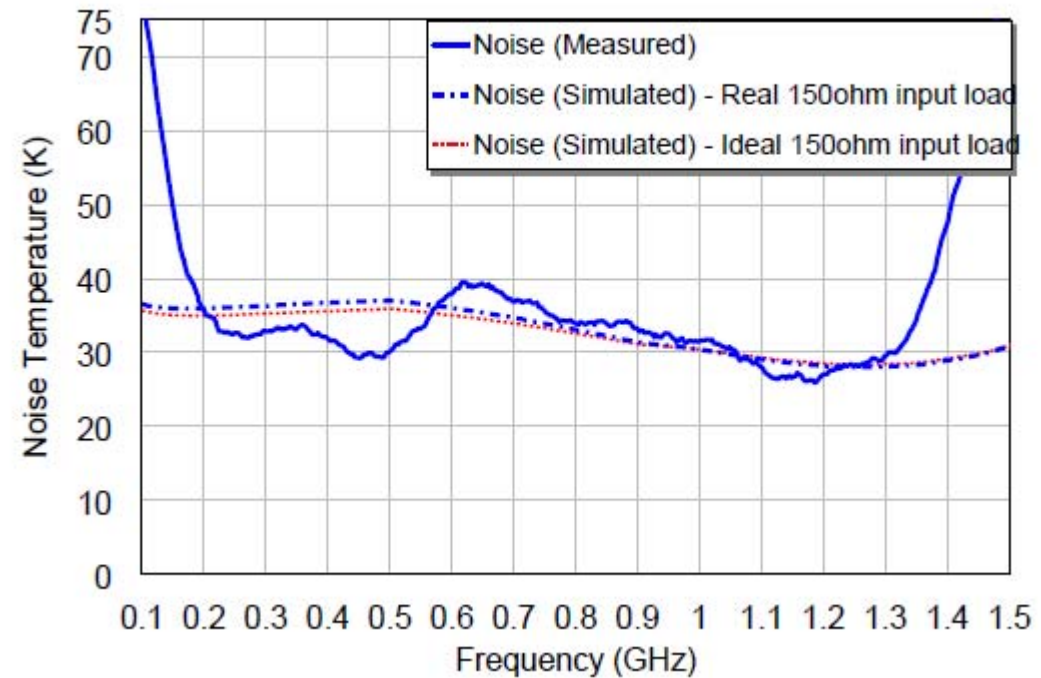
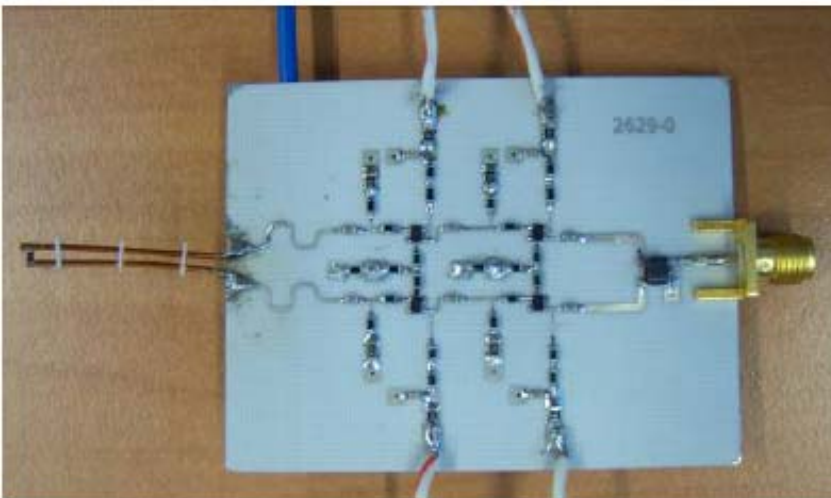
Simulations Ref. Mateos et.al., 2004/2008

- And analyses based on a normalization of Γ_{act} to S_{11} of the LNA: reasonable power match required!



Avago 35143 differential LNA

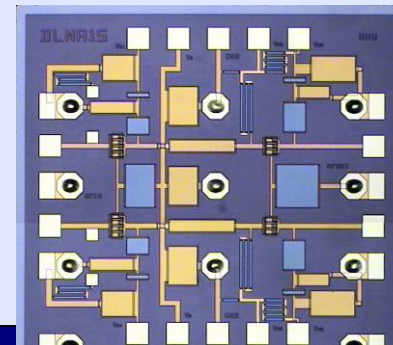
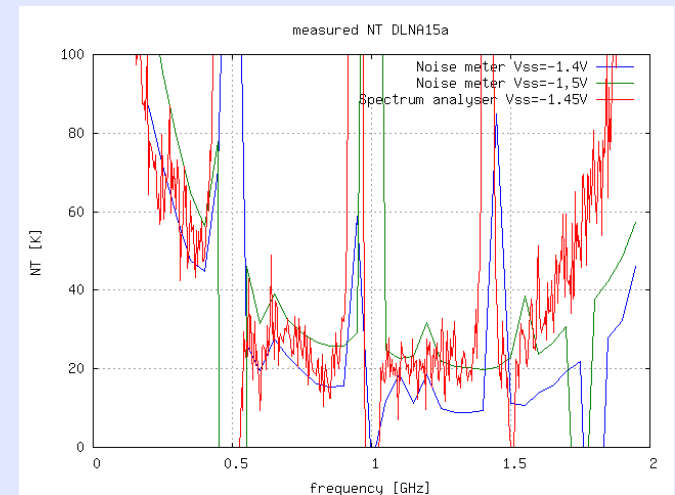
- 150 Ohm input impedance





MMIC LNA

- mHEMT OMMIC LNAs
 - differential
 - Low R_n , Low Noise @300K
 - $f_T \sim 250\text{GHz}$
 - Measurement $\sim 20\text{K}$
 - Oscillations issues

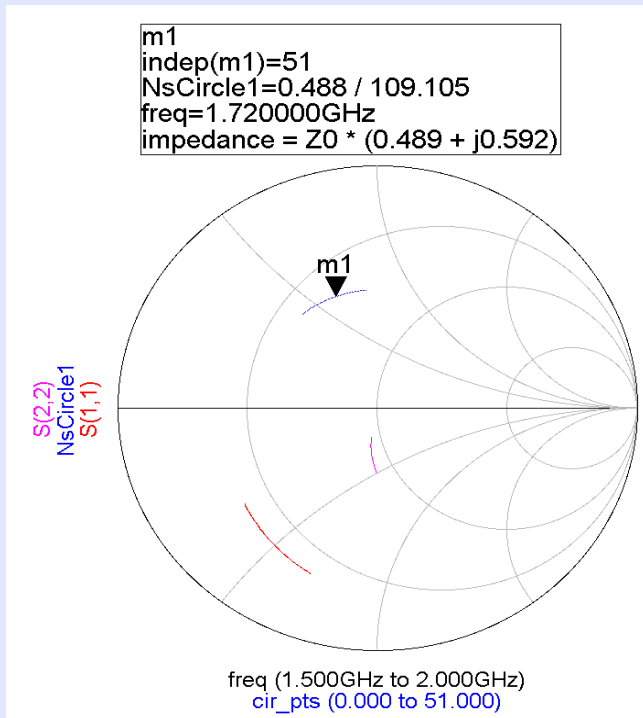




70nm mHEMT process validation

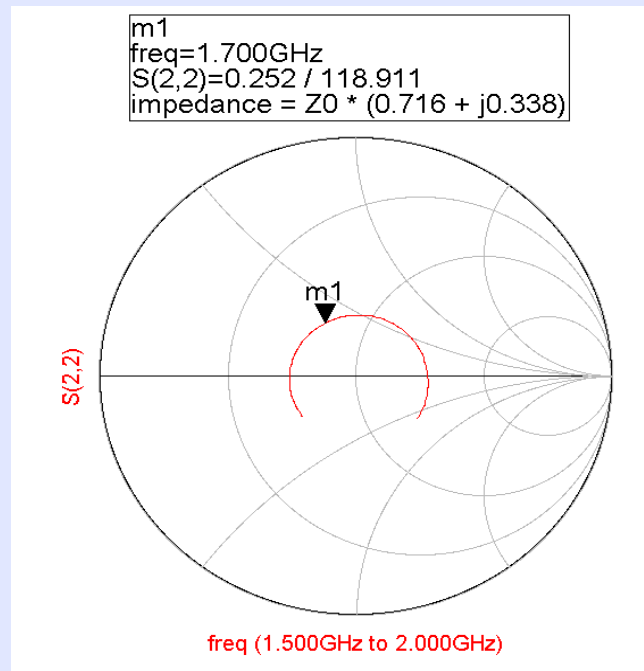


To validate the noise modelling of the mHEMT, the Γ_{opt} and the R_n have been measured: models not correct



simulation

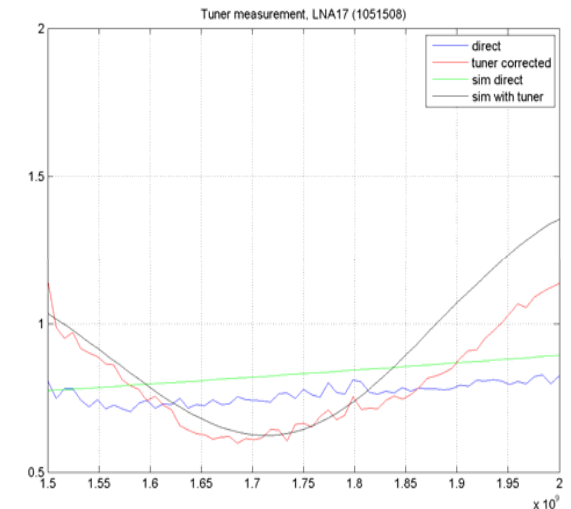
$R_n=0.17$



measured

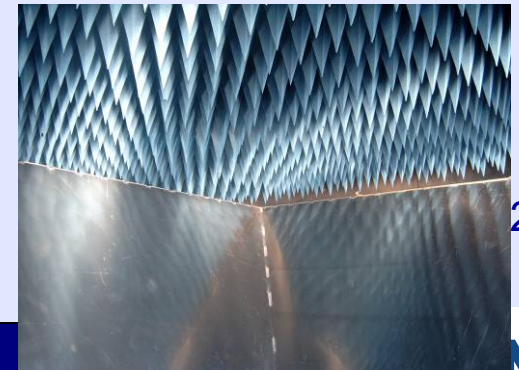
$R_n=0.19$

@1.6GHz





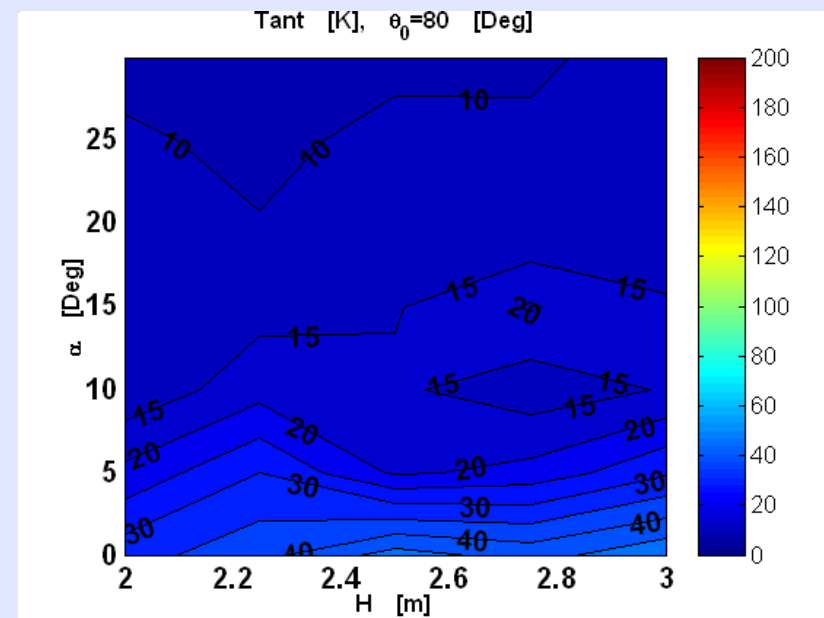
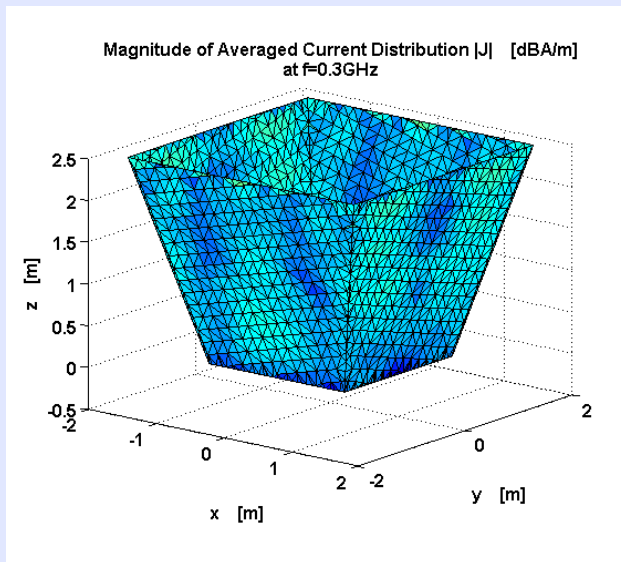
Hot cold test facility: THACO





THACO analyses

- Spill over simulations

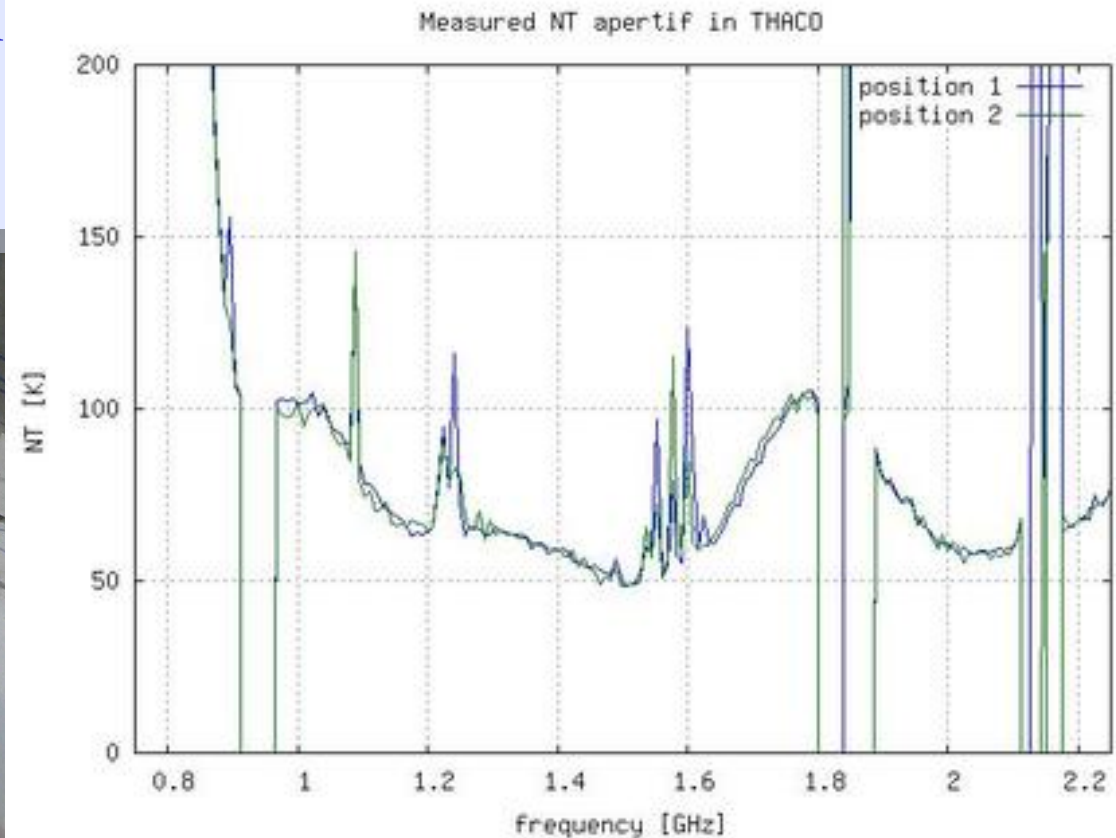
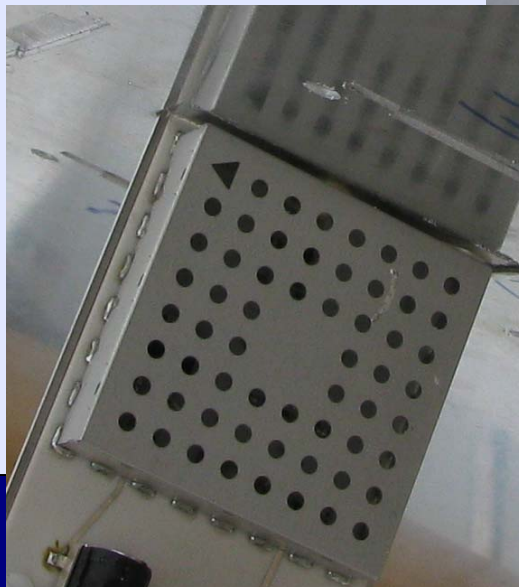




Antenna: Al Vivaldi radiator

PAFs; APERTIF

- Feed board and LNA integrated on one board: Avago based APERTIF Vivaldi+LNA
 - Measured 52K Trec







Pitfalls

- Unwanted visitors



✓ More RFI



6



AA T_{sys} characterization

- Action plan..
 - Establish available equipment list
 - Including noise parameter test sets, tuner set-up / F50 method
 - Select AA relevant LNA for cross testing: *Transfer LNA*
 - Including approach for differential LNAs
 - Set-up small working group → frequent cross test of new designs
- Involvement off all partners working on low noise
 - UoM, OPAR, FG-IGN, Chalmers, Calgary, ASTRON.....and?
 - Start with SKADS soon evolving in *Aperture Array Verification Plan*



Conclusion

- More effort required on room temperature LNAs
- Low frequency transistor modeling:
 - Pospieszalski, many unphysical models
- Low frequency array T_{sys} characterization
- Design of a 1500m² prototype array starts soon
 - 100.000 LNAs