



Shanghai Astronomical Observatory
Chinese Academy of Sciences

Measuring RFI data loss

Willem Baan

SHAO & Linnæus Univ & ASTRON

support from Axel Jessner
and Netherlands Administration

ITU-R Regulatory Conditions

- RR 5.340 - All emissions are prohibited in selected bands
bands like 1427 MHz
- RR 5.149 - Administration take all practical steps to protect
operations of the Radio Astronomy Service in the
following bands:
- ITU-R RA.769-2 - these are the thresholds above which
interference is detrimental for spectral line and
continuum observations
- ITU-R RA.1513 - in primary bands of the RAS 2% data loss in
time is allowed for single system and 5% for the
aggregate

Practical Thresholds (from RA.769-2)

- ITU-R RA.769-2 puts into words variation of radiometer noise of a standard RAS system with $T = T_A + T_R$

$$\Delta P/P = \Delta T/T = 1/\sqrt{\Delta t \Delta f}$$

- Assuming a harmful threshold of 10% error of radiometric power:

$$\begin{aligned}\Delta P_H &= 0.1 \Delta P \Delta f = 0.1 k \Delta T \Delta f \\ &= 0.1 k T(\text{sys}) \Delta t^{-1/2} \Delta f^{1/2}\end{aligned}$$

- Adopting values for T_{sys} and reference values for bandwidth Δf and for integration time Δt of 2000 sec

- Harmful pfd: $S_H \Delta f = \Delta P_H + 10 \log f(\text{Hz}) - 158.5 \quad (\text{dB(W/m}^2\text{)})$
- Harmful spfd: $S_H = \Delta P_H + 5 \log f(\text{Hz}) - 158.5 \quad (\text{dB(W/m}^2\text{/Hz)})$

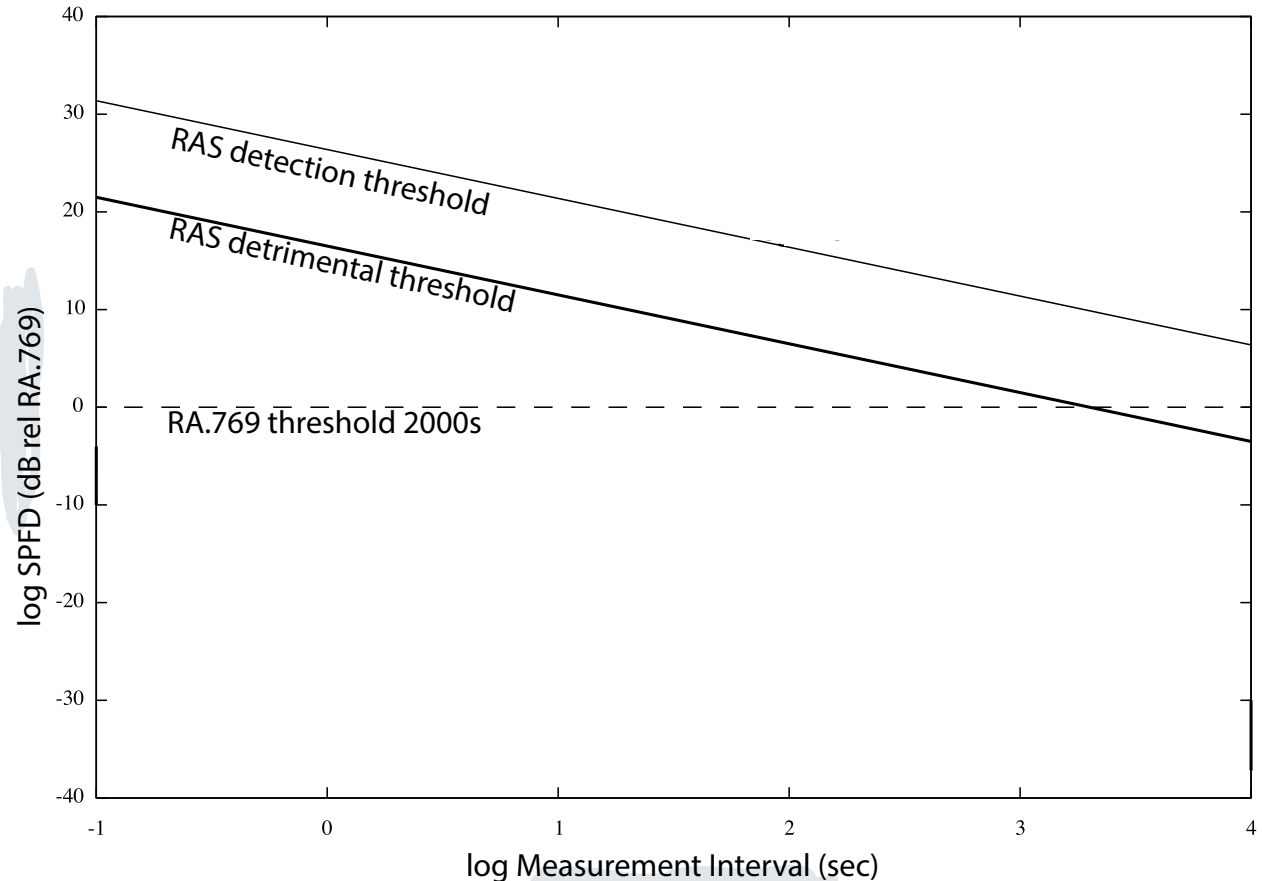
RA.769-2 thresholds for RAS antenna

$$T_{\text{spec}}(\Delta t, \Delta f) = \text{spfd}_{\text{spec}}(\text{RA.769, table 2}) + 5 \log \left(\left(\frac{\Delta f}{\Delta f_{\text{ref}}} \right) \left(\frac{2000}{\Delta t} \right) \right) \\ (\text{dB(W/m}^2\text{/Hz)})$$

$$T_{\text{cont}}(\Delta t, \Delta f) = \text{spfd}_{\text{cont}}(\text{RA.769, table 1}) + 5 \log \left(\left(\frac{\Delta f}{\Delta f_{\text{b}}} \right) \left(\frac{2000}{\Delta t} \right) \right) \\ (\text{dB(W/m}^2\text{/Hz)})$$

spfd for RA.769-2
varies w measurement
 Δf and Δt

Time interval of
2000s and frequency
channel of 20 kHz are
reference values
=> NOT 'holy numbers'



The Measurement System

- Detector system with forward gain G and equal/inferior T_{sys}
- Sensitivity dependence to measurement Δf and Δt remains same
- System G gives vertical scaling of detection noise floor

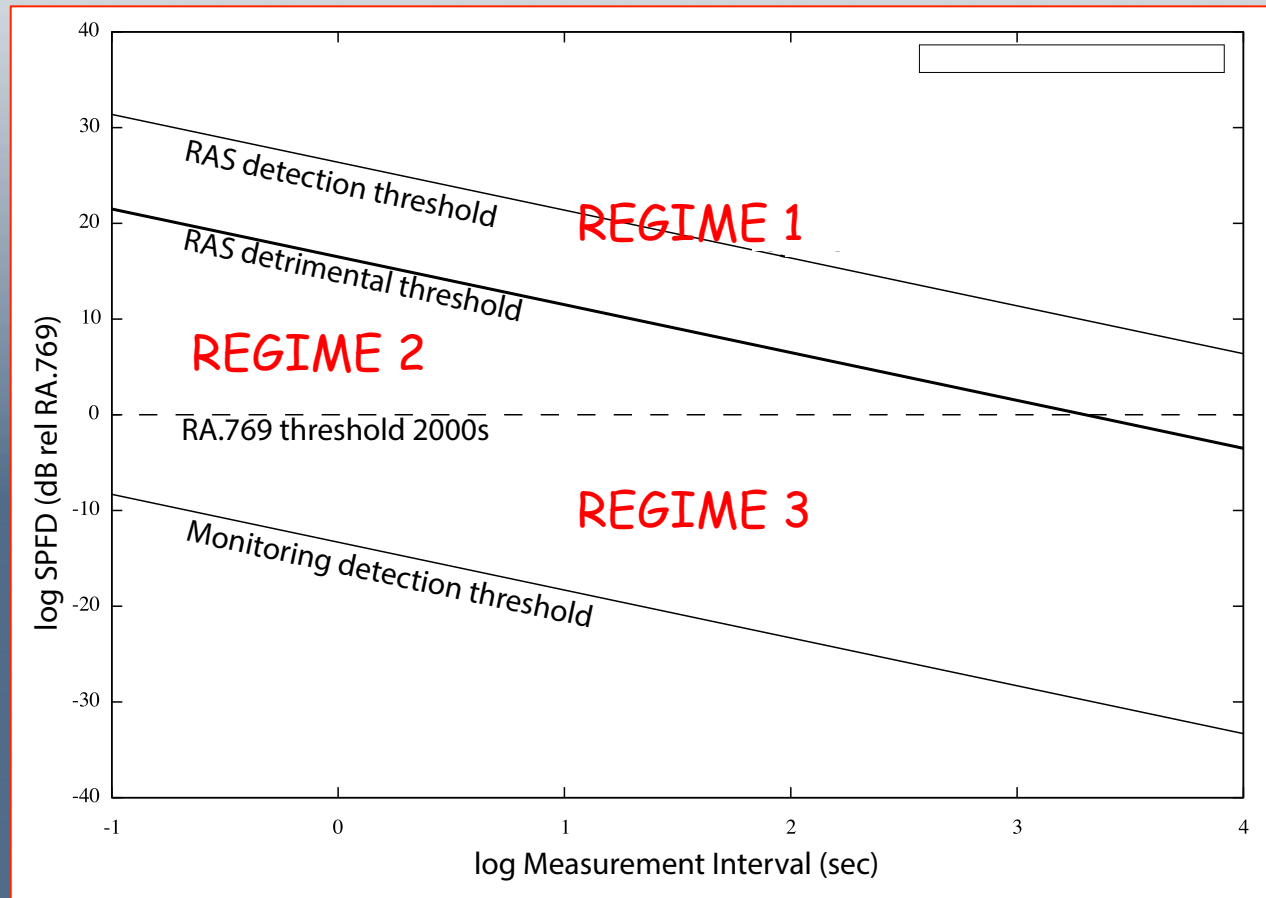
$$S_{\text{spec}}(\Delta t, \Delta f) = - (G+10) + [10 \log (T_{\text{sys,ref}}/T_{\text{sys,mon}})] + T_{\text{spec}}(\Delta t, \Delta f) \\ (\text{dB(W/m}^2\text{/Hz)})$$

$$S_{\text{cont}}(\Delta t, \Delta f) = - (G+10) + [10 \log (T_{\text{sys,ref}}/T_{\text{sys,mon}})] + T_{\text{cont}}(\Delta t, \Delta f) \\ (\text{dB(W/m}^2\text{/Hz)})$$

The Measurement Antenna

Antenna lowers the
detection threshold

Three distinct RFI
regimes



Methodology

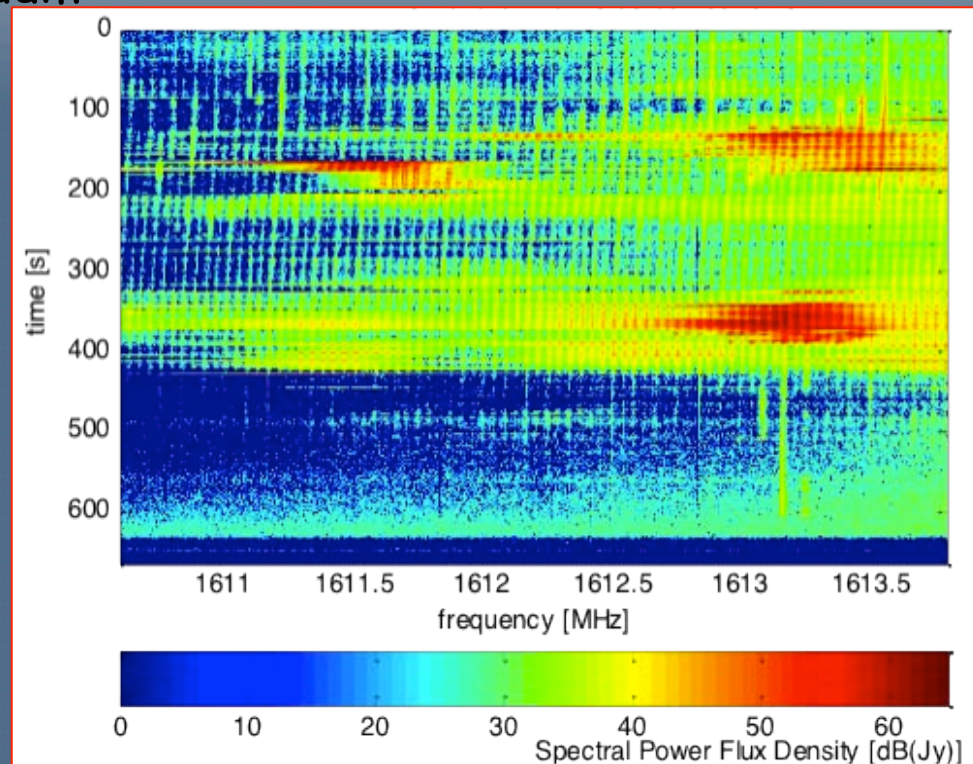
- Calibrate the measurement system
 - \Rightarrow calibrate T_A conversion to Jy or dB(W/m²/Hz)
 - \Rightarrow example: ON-OFF on source with known flux density (Cas A or Cen A) and find Jy/K ratio
- Identify interference characteristics (variability in time & freq)
- Take time-series with taking small Δt and appropriate Δf
 - \Rightarrow longer than characteristic time/freq variability timescale
 - \Rightarrow choice of Δt and Δf affects the percentage data loss
- Determine percentage of time where RFI exceeds the corrected RA.769 threshold levels
- Determine integrated RFI over long time interval
 - \Rightarrow for primary bands and RR 5.340 bands
 - $\Rightarrow N\Delta t$ long enough to cover variability (e.g. 2000s)

Evaluation of Results

- Primary bands - is the percentage of time larger than corrected RA.769 threshold more than designated by RA.1513 ?
=> separate evaluation for continuum and spectral line
- RR 5.340 bands - in addition: is the integrated RFI larger than appropriate values from RA.769 ?
=> spectral line & continuum

Time-frequency occupancy diagram of unwanted emissions in a single passage of a LEO MSS satellite. Diagram contains $N = 630$ data records with $d\Delta t = 1$ s and $M = 420$ spectral channels with $d\Delta f = 6.1$ kHz within the RAS band. The percentage varies from about 30% to 100%

ECC Report (09)02
(Leeheim - BNetzA and MPIfR)



Determining Percentages

- **Time series analysis**
- P = number of affected records from total N records
- \Rightarrow percentage = $P/N \times 100\%$ valid for large N
- $\Rightarrow N$ needs to be large for small percentages
- For relatively small N and unpredictable RFI
- \Rightarrow percentage = $(P + 1)/(N + 2) \times 100\%$
- \Rightarrow probability of random RFI to enter into P out of N

- **Time-occupancy analysis**
- P = number of affected $(\Delta t, \Delta f)$ elements in $N \times M$ pixels
- \Rightarrow percentage = $(P / N * M) \times 100\%$
- \Rightarrow regridding of frequency axis conform ITU-R RA.769
- \Rightarrow coarser Δt & Δf raise the percentage data loss

The Impact of RFI on data

- Straight forward measurements do not reflect the impact
- Impact of RFI is different for spectral line and continuum data
- Time and frequency variability ?
- Location of RFI inside spectrum important for spectral line but not for continuum
- This is not discussion (yet) on translation of percentage data loss into actual data loss
 - => how much can we flag ?
 - => what is really lost with 2% and 5% in time ?
 - => collateral damage from weaker sidebands & spurious ?

ITU-R action

- Procedures outlined here are submitted as document 7D/48-E (2013) to ITU-R WP7D
- Revised version of the PDNR by NL & GER